

# Lecture 3 --Review of Quantum Physics & Basic AMO Physics

Version #1 -- 2/9/2014

## I Some important model systems in QM

- i) 2-state QM ( $d=2$ , simplest hilbert space)
  - spin-1/2 ( $\uparrow$  &  $\downarrow$ )
  - photon 2-polarization ( $\leftrightarrow$  &  $\hat{\uparrow}$ )
- ii) Particle in box /1D Schrodinger
- iii) Harmonic oscillator
- iv) Hydrogen atom  $\rightarrow$  AMO physics
- v) Periodic potential  $\rightarrow$  solid state physics

## II Foundations of QM: early milestones

- i) Photon (Planck blackbody radiation)
  - $E=hf=\hbar\omega$

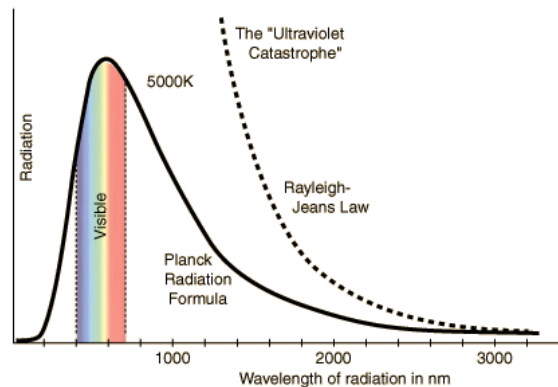


Fig 3.1

Total energy per unit volume in the cavity in  $\nu$  and  $\nu+d\nu$

Rayleigh – Jeans formula:  $u(\nu)d\nu = \varepsilon G(\nu)d\nu = (8\pi kT/c^3)\nu^2 d\nu$

Plank radiation formula:  $u(\nu)d\nu = (8\pi h/c^3)(\nu^2 d\nu)/(e^{h\nu/kT} - 1)$

- ii) Photo-electric effect & Compton scattering (light matter interaction)
  - ... *PE can measure Planck constant*
  - Setup of Photo-electric effect experiment: setup and Experiment Results

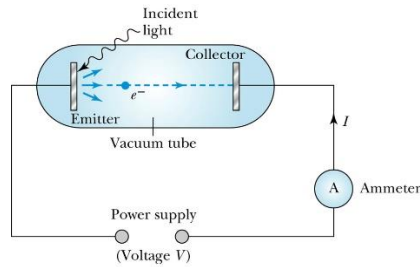


Fig. 3.2 experiment setup

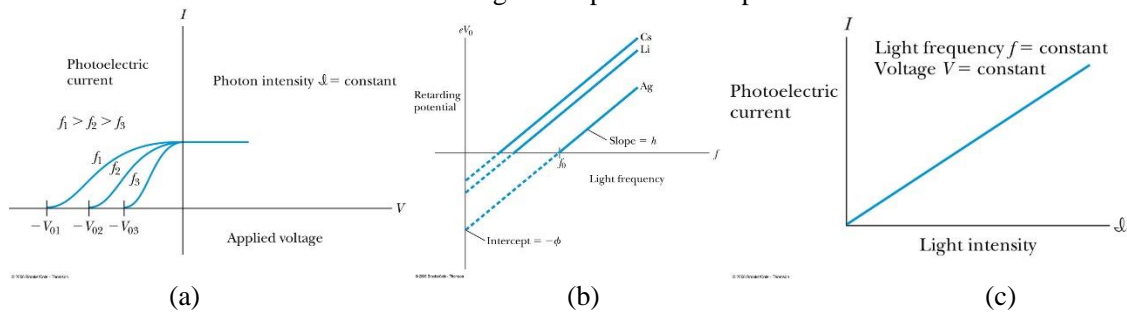


Fig. 3.3 Experiment results

### Einstein's Theory

- Conservation of energy yields:  
Energy before (photon) = energy after (electron)  
$$hf = \phi + \text{K.E. (electron)}$$
where  $\phi$  is the work function of the metal.  
Explicitly the energy is  $hf = \phi + \frac{1}{2}mv_{\text{max}}^2$
- The retarding potentials measured in the photoelectric effect are the opposing potentials needed to stop the most energetic electrons.

$$eV_0 = \frac{1}{2}mv_{\text{max}}^2$$

### Quantum Interpretation

- The kinetic energy of the electron does not depend on the light intensity at all, but only on the light frequency and the work function of the material.

$$\frac{1}{2}mv_{\text{max}}^2 = eV_0 = hf - \phi$$

- Einstein in 1905 predicted that the stopping potential was linearly proportional to the light frequency, with a slope  $h$ , the same constant found by Planck. (fig. 3.3 (b))

$$eV_0 = \frac{1}{2}mv_{\text{max}}^2 = hf - hf_0 = h(f - f_0)$$

- From this, Einstein concluded that light is a particle with energy:

$$E = hf = \frac{hc}{\lambda}$$

iii) Spectroscopy (atomic physics/astronomy) --- start from hydrogen --- Bohr's model (atom/matter quantized levels)

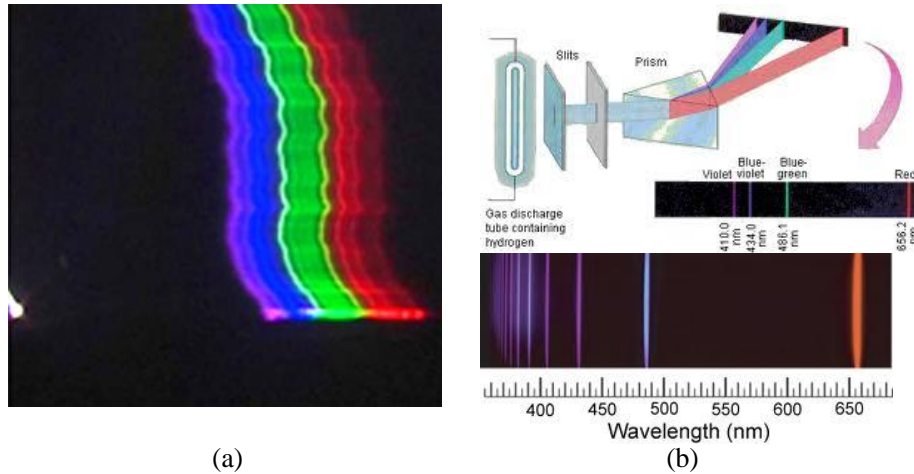


Fig. 3.4 (a) Spectrum of Ball Lightning (b) Hydrogen Spectrum

**External link:**

<http://physics.aps.org/articles/v7/5>

Observation of the Optical and Spectral Characteristics of Ball Lightning

Jianyong Cen, Ping Yuan, and Simin Xue

[Phys. Rev. Lett. \*\*112\*\*, 035001 \(2014\)](#)

### III Concepts & theory

i) Matter wave; wave-particle duality (De Broglie)

$$\text{Energy: } E = mc^2 = \frac{m_0 c^2}{\sqrt{1 - \frac{v^2}{c^2}}} = \sqrt{m_0^2 c^4 + P^2 c^2} = \sqrt{E_0^2 + P^2 c^2}$$

$m_0 = \text{rest mass}$

$E_0 = \text{rest energy}$

$$\text{Momentum: } P = \frac{m_0 v}{\sqrt{1 - \frac{v^2}{c^2}}} = h/\lambda = h/(mv)$$

**Comparing to Photon (massless particle):**  $E = hf = pc \Rightarrow p = hf/c = h/\lambda$

**$P = h/\lambda$  is true for all particles**

Wave Optics: diffraction

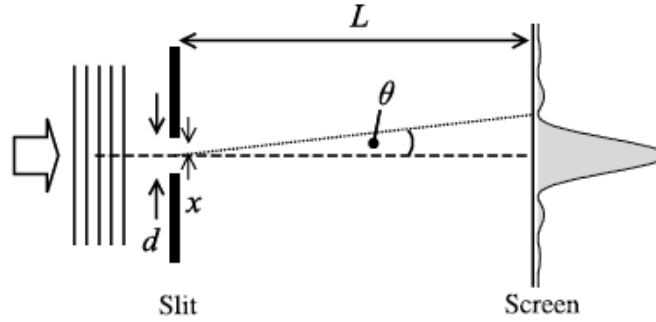


Fig. 3.5 When the distance  $L$  between the slit and screen is larger than the **Rayleigh Distance** ( $d^2/\lambda$ ), the diffraction pattern is said to be in the **far-field (Fraunhofer) limit**. On the other hand, when  $L \lesssim d^2/\lambda$ , we are in the **near-field (Fresnel) regime**. In what follows, we consider only Fraunhofer limit

$$\text{Diffraction pattern: } I(\theta) \propto \left(\frac{\sin\beta}{\beta}\right)^2$$

$$\text{Resolution: } \sin\theta_{\min} = \pm \frac{\lambda}{d}$$

$$\text{resolution for circular aperture: } \sin\theta_{\min} = 1.22 \frac{\lambda}{d}$$

ii) Uncertainty principle (Heisenberg)

– Fourier transform picture

$$\text{Heisenberg Uncertainty Principle: } \Delta x \Delta p_x \geq \hbar/2$$

## IV Theory & Formulation

i) Schrodinger equation/wavefunction

$$\mathcal{P}(\mathbf{r}, t) dV = |\Psi(\mathbf{r}, t)|^2 dV. \quad \text{-- probability density}$$

The equation of motion of the wave function is given by the **Schrödinger equation**:

$$\hat{H}\Psi(\mathbf{r}, t) = i\hbar \frac{\partial}{\partial t} \Psi(\mathbf{r}, t), \quad (3.2)$$

$$\hat{H} = \frac{\hat{p}^2}{2m} + \hat{V}(\mathbf{r}) = -\frac{\hbar^2}{2m} \nabla^2 + \hat{V}(\mathbf{r}). \quad \text{“Hamiltonian”}$$

The position operator is given by

$$\hat{\mathbf{r}} = \mathbf{r},$$

the momentum operator is:

$$\hat{p} = -i\hbar\nabla.$$

$$\Psi(\mathbf{r}, t) = \psi(\mathbf{r}) \Theta(t). \quad \Psi(\mathbf{r}, t) = \psi(\mathbf{r}) \exp(-iEt/\hbar).$$

time-independent Schrödinger equation:

$$\hat{H}\psi(\mathbf{r}) \equiv -\frac{\hbar^2}{2m}\nabla^2\psi(\mathbf{r}) + \hat{V}(\mathbf{r})\psi(\mathbf{r}) = E\psi(\mathbf{r}).$$

### Several Should Know Points

- Example: free particle:  $\hat{H}\psi_n(\mathbf{r}) = E_n\psi_n(\mathbf{r}),$
- Probability density:  $\int_{-\infty}^{+\infty} |\psi(x)|^2 dx = 1,$
- Normalizable (bound) vs unnormalizable (extended) state
- Superposition state  $\psi = \sum_i c_i \varphi_i$
- Measurement(!..“weak”measure) & expectation value/ variance  

$$(\Delta O)^2 = \int \psi^* (\hat{O} - \langle \hat{O} \rangle)^2 \psi d^3r$$

$$= \langle \hat{O}^2 \rangle - \langle \hat{O} \rangle^2.$$
- Commutator/commutation & general uncertainty principle  

$$[\hat{A}, \hat{B}] \equiv \hat{A}\hat{B} - \hat{B}\hat{A}. \quad \Delta A)^2(\Delta B)^2 \geq \left| [\hat{A}, \hat{B}] \right|^2 / 4,$$

$$[\hat{x}, \hat{p}_x] = i\hbar. \quad \Delta x \Delta p_x \geq \hbar/2.$$

ii) Matrix formulation (Heisenberg)

iii) Dirac “bra-ket” ... quantum state/Hilbert space

“Hamiltonian” [*“Abrahams example”*]  $\hat{H} = \hat{T} + \hat{V},$

Wave function <-> vector in Hilbert space

### Particle in a box/1D Schrodinger equation

- Infinite square well (uncertainty)
- Finite square well
- 1D scattering state (refection & transmission --- “quantum reflection” & tunneling, even “resonant tunneling”)

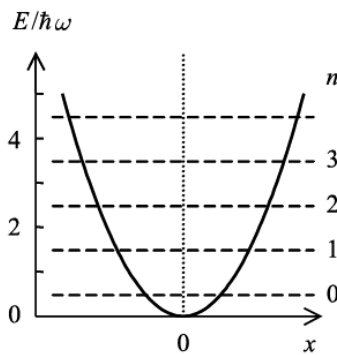
### Harmonic oscillator

- Harmonic trap
- Quantized EM fields...

$$-\frac{\hbar^2}{2m} \frac{d^2\psi(x)}{dx^2} + \frac{1}{2}m\omega^2 x^2 \psi(x) = E\psi(x).$$

$$\psi_n(x) = u_n(x) \exp(-m\omega x^2/2\hbar),$$

$$E_n = \left(n + \frac{1}{2}\right) \hbar\omega,$$



$n$	$u_n(x)$
0	$(1/a\sqrt{\pi})^{1/2}$
1	$(1/2a\sqrt{\pi})^{1/2} 2(x/a)$
2	$(1/8a\sqrt{\pi})^{1/2} [2 - 4(x/a)^2]$
3	$(1/48a\sqrt{\pi})^{1/2} [12(x/a) - 8(x/a)^3]$
$\vdots$	
$n$	$(1/n!2^n a\sqrt{\pi})^{1/2} H_n(x/a)$

Harmonic length  $a = (\hbar/m\omega)^{1/2}$

for the  $n$ th level (see Exercise 3.11):

$$\Delta x \Delta p_x = \left(n + \frac{1}{2}\right) \hbar.$$

*Lecture3 notes are collected by Bohao Liu. Correction, supplement and suggestion are greatly welcomed. Bohao Liu's email is liubohao@purdue.edu.*